

Fermi-Dirac Distribution : Derivation

start with probability weighting factors using the canonical ensemble

$$P_N(E) = \frac{e^{-E/h_bT}}{\sum e^{-E_\alpha^N/h_bT}}$$

E_α^N is a stationary state of the system

$$\sum e^{-E_\alpha^N/h_bT} = \text{partition function}$$

related to Helmholtz free energy

$$F = U - TS$$

U = internal energy ; S = entropy

$$\sum e^{-E_\alpha^N/h_bT} = e^{-F_N/h_bT}$$

$$\Rightarrow P_N(E) = e^{-(E-F_N)/h_bT}$$

$$\text{define } f_i^N = \sum P_N(E_\alpha^N)$$

From the above and after a short derivation, it is straightforward to show that:

$$f_i = \frac{1}{e^{(\epsilon_i - \mu)/k_B T} + 1}$$

f_i is the mean number of electrons in the i -electron level ϵ_i

total number of electrons $N = \sum_i f_i$

$$\Rightarrow N = \sum_i \frac{1}{e^{(\epsilon_i - \mu)/k_B T} + 1}$$

From now, we will call f the Fermi function

$$f(\epsilon) = \frac{1}{e^{(\epsilon - \mu)/k_B T} + 1}$$

Note that $f(\epsilon)$ gives the ground state distribution at $T = 0$

$$\left. \begin{aligned} \lim_{T \rightarrow 0} f_{k_B} &= 1 & \epsilon(\vec{k}) < \mu \\ &= 0 & \epsilon(\vec{k}) > \mu \end{aligned} \right\} \begin{array}{l} \text{consistent} \\ \text{with} \\ \text{ground} \\ \text{state} \end{array}$$

$$\Rightarrow \lim_{T \rightarrow 0} \mu = \epsilon_F$$

Towards thermal properties of metals
using Fermi function $f(\epsilon)$

constant-volume specific heat

$$\cancel{C_v = \frac{1}{V} \left(\frac{\partial S}{\partial T} \right)_V} \quad u = \frac{U}{V} = \frac{F + TS}{V}$$

$$C_v = \left(\frac{\partial u}{\partial T} \right)_V = \left(\frac{\partial u}{\partial S} \right)_V \left(\frac{\partial S}{\partial T} \right)_V = \frac{1}{V} \left(\frac{\partial S}{\partial T} \right)_V$$

$$C_v = \left(\frac{\partial u}{\partial T} \right)_V = \frac{1}{V} \left(\frac{\partial S}{\partial T} \right)_V$$

$$U = 2 \sum_{\vec{k}} \epsilon(\vec{k}) f(\epsilon(\vec{k}))$$

$$\rightarrow u = \frac{U}{V} = \frac{2}{V} \sum_{\vec{k}} \epsilon(\vec{k}) f(\epsilon(\vec{k}))$$

$$u = \int \frac{d\vec{k}}{4\pi^3} \epsilon(\vec{k}) f(\epsilon(\vec{k}))$$

recall that

$$\Delta \vec{k} = \frac{8\pi^3}{V}$$

$$\frac{1}{\Delta \vec{k}} = \frac{V}{8\pi^3}$$

Similarly

$$n = \int \frac{d\vec{k}}{4\pi^3} f(\epsilon(\vec{k}))$$

General Function $F(\vec{r})$

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$$\int \frac{d\vec{h}}{4\pi^3} F(\varepsilon(\vec{h})) = \int_{-\infty}^{\infty} d\varepsilon g(\varepsilon) F(\varepsilon)$$

$$\left. \begin{aligned} g(\varepsilon) &= \frac{m}{\hbar^2 \pi^2} \sqrt{\frac{2m\varepsilon}{\hbar^2}}, \quad \varepsilon > 0 \\ g(\varepsilon) &= 0, \quad \varepsilon < 0 \end{aligned} \right\} \text{density of levels}$$

which can also be written as:

$$g(\varepsilon) = \frac{3}{2} \frac{n}{\varepsilon_F} \left(\frac{\varepsilon}{\varepsilon_F} \right)^{1/2}, \quad \varepsilon > 0$$

$$g(\varepsilon) = 0, \quad \varepsilon < 0$$

Then, at the Fermi level,

$$g(\varepsilon_F) = \frac{m k_F}{\hbar^2 \pi^2}$$

or

$$g(\varepsilon_F) = \frac{3}{2} \frac{n}{\varepsilon_F}$$

Using the density of levels $g(\epsilon)$,

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$$u = \int_{-\infty}^{\infty} d\epsilon g(\epsilon) \epsilon f(\epsilon)$$

and

$$n = \int_{-\infty}^{\infty} d\epsilon g(\epsilon) f(\epsilon)$$

} energy density
and
number density
in
terms of
 $\epsilon, g(\epsilon), f(\epsilon)$

Evaluating above equations is not simple

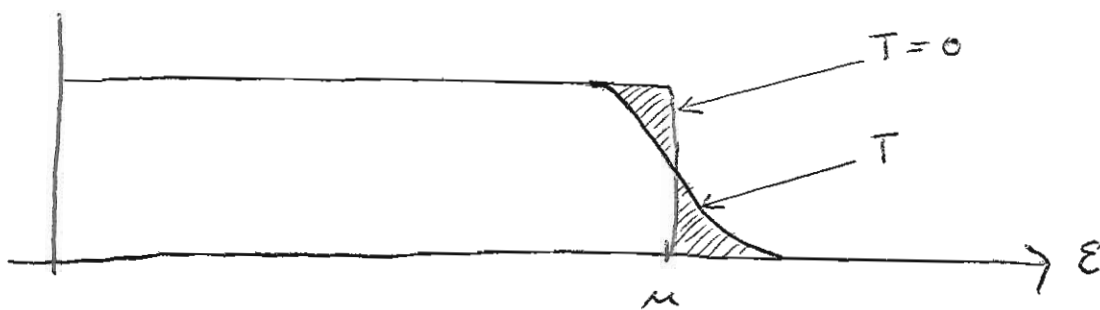
But fortunately, there exists a systematic expansion which can be applied for most temperatures of interest

- makes use of fact that $k_B T \ll k_B T_F$

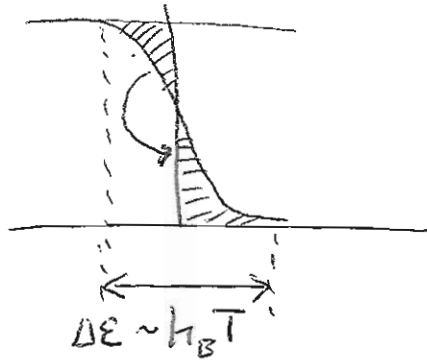
at room temperature,

$$\frac{k_B T}{k_B T_F} \sim \frac{3 \times 10^2 \text{ K}}{3 \times 10^4 \text{ K}} = \underline{.01}$$

→ means that



First, though, consider how much thermal energy density is excited when going from $T=0$ to $T=T$



number density of excited electrons is:

$$n_{\text{ex}} \approx k_B T g(\epsilon_F)$$

excitation energy/electron is

$$\epsilon_{\text{ex}} \approx k_B T$$

$$\Rightarrow \Delta u \approx n_{\text{ex}} \epsilon_{\text{ex}} = g(\epsilon_F) (k_B T)^2$$

$$c_v = \left(\frac{du}{dT} \right)_n = \left[2 g(\epsilon_F) k_B^2 \right] T$$

if we had another factor of $\frac{\pi^2}{6} \sim \frac{3}{2}$

we would already get the correct result

$$c_v = \left[\frac{\pi^2}{3} g(\epsilon_F) k_B^2 \right] T$$

$$\text{But } g(\epsilon) = \frac{3}{2} \frac{n}{\epsilon_F}$$

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$$\Rightarrow C_v = \frac{\pi^2}{2} \frac{n}{\epsilon_F} h_B^2 T$$

$$C_v = \frac{\pi^2}{2} \left(\frac{h_B T}{\epsilon_F} \right) n h_B$$

compare to classical

$$C_v^{\text{class}} = \left[\frac{3}{2} \right] n k_B$$

$$\Rightarrow C_v = A C_v^{\text{class}}$$

$$A = \frac{\pi^2}{3} \left(\frac{h_B T}{\epsilon_F} \right)$$

$$A \sim .01 = 10^{-2}$$

$\Rightarrow C_v$ smaller than classical one
by 10^{-2}

Note: this is part of what ~~Drude theory missed~~

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